

Improving the Mass Reach

Light-Shining-Through-Walls Experiments with Phase Shift Plates

Joerg Jaeckel¹
Andreas Ringwald²

arXiv:0706.0693 [hep-ph]

¹IPPP Durham, ²DESY Hamburg

Introduction

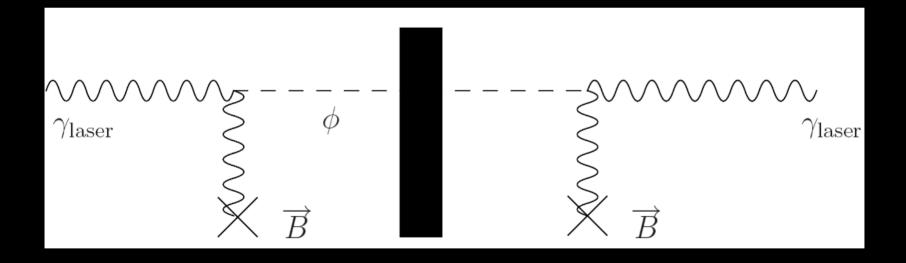
LSW Experiments

Light Shining Through Walls: ALPs



Interaction

$$\mathcal{L} \sim g\phi F^{\mu\nu}\tilde{F}_{\mu\nu} \sim g\phi \vec{E} \cdot \vec{B}$$

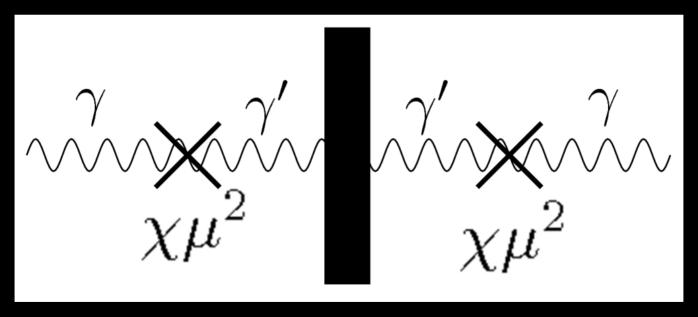


Light Shining Through Walls: Paraphotons



Two U(1)'s + with mixing term

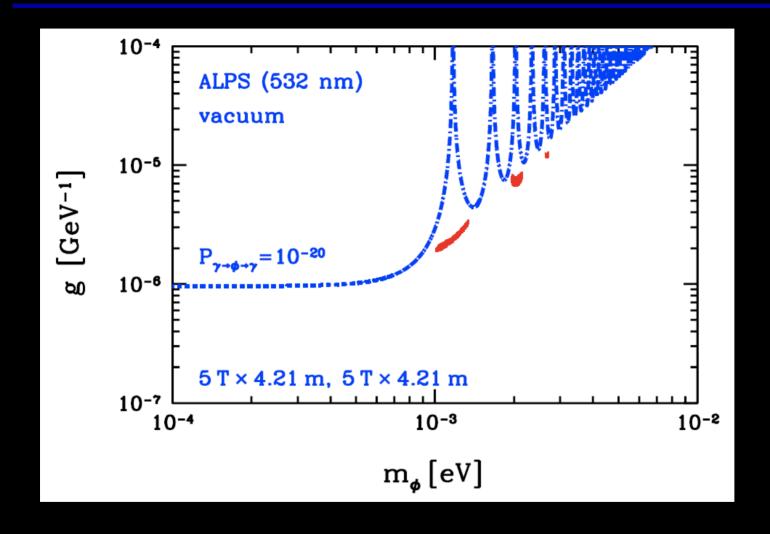
$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} - \frac{1}{4}B^{\mu\nu}B_{\mu\nu} + \frac{1}{2}\mu^2(B^{\mu}B_{\mu} - 2\chi A^{\mu}B_{\mu} + \chi^2 A^{\mu}A_{\mu})$$



Problem for higher masses

Problem

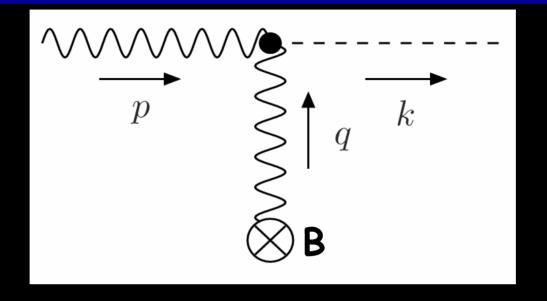






Reason: Version I





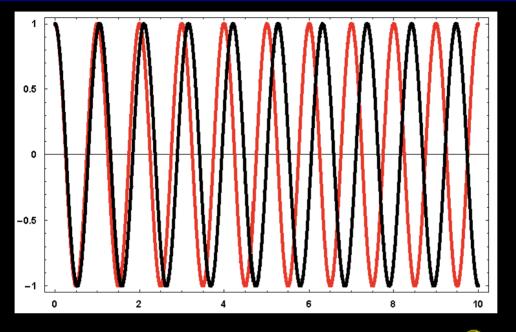
$$|\vec{q}| = |\vec{p}| - |\vec{k}| = \omega - \sqrt{\omega^2 - m^2} \approx \frac{m^2}{2\omega} > 0$$

Need inhomogeneous magnetic field with

$$\frac{1}{L} \gtrsim |\vec{q}|$$

Reason: Version II





$$\frac{2\pi}{\omega} = \lambda_{\gamma} < \lambda_{\text{ALP}} = \frac{2\pi}{\sqrt{\omega^2 - m^2}}$$

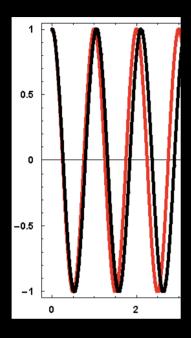
ALP and photon field run out of phase



Solution

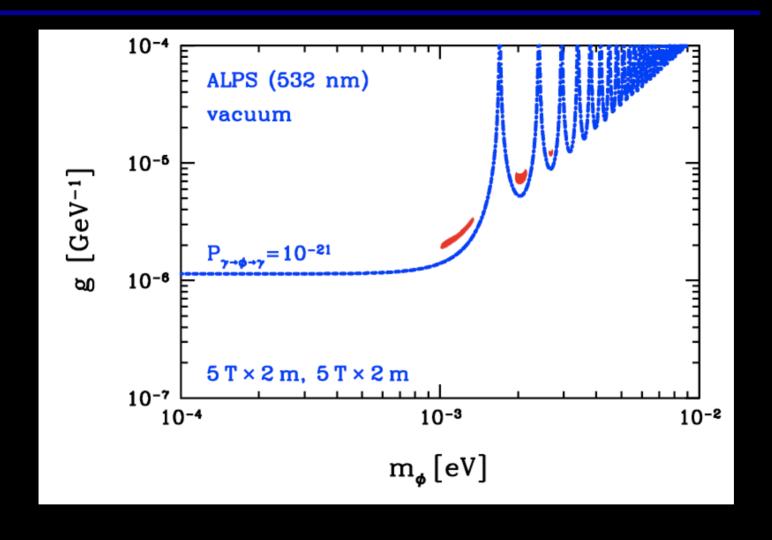
Idea 0: Cut Magnet





Idea 0: Problem







Idea I: Alternating Magnetic Field





Inhomogeneities of size $\sim \frac{L}{N}$

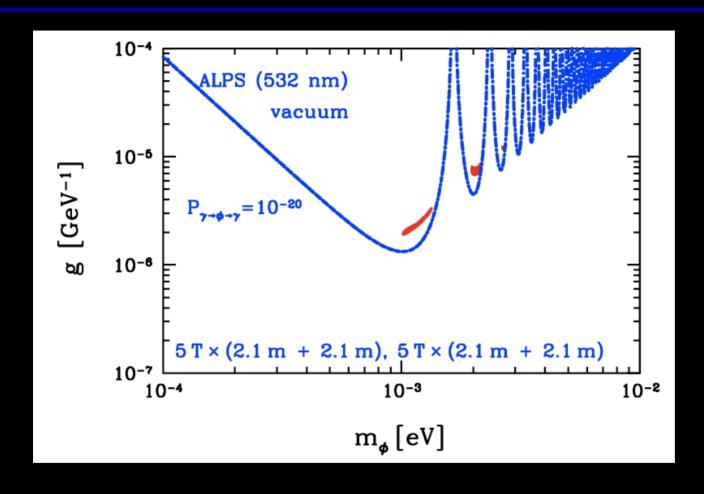
magnetic field can supply N times more momentum



Can reach higher masses!

Improvement I







But, reduced sensitivity at small masses!

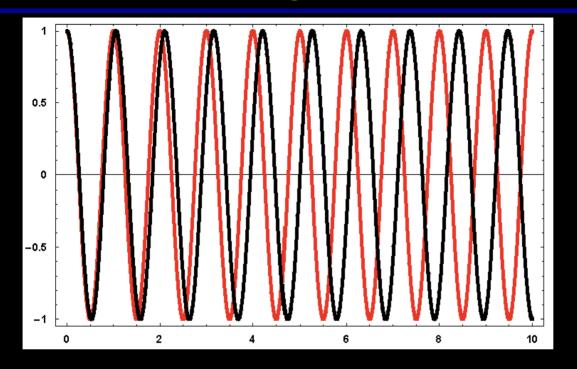
Difficulties I



- Need more than one magnet.
- · Sensitivity at small masses is reduced.
- Difficult to change the # of segments.

Idea IIa: Gas Filling







Wavelength of Photon and ALP differ

We can modify the wavelegth of the photon!

$$\lambda = rac{\lambda_{ ext{vac}}}{n}$$

Difficulties IIa



$$\frac{2\pi}{\omega} = \lambda_{\gamma} < \lambda_{\text{ALP}} = \frac{2\pi}{\sqrt{\omega^2 - m^2}}$$

$$\lambda_{\gamma} = rac{\lambda_{ ext{vac}}}{n}$$
 \longrightarrow We need n<1!

Difficult for low energy photons!

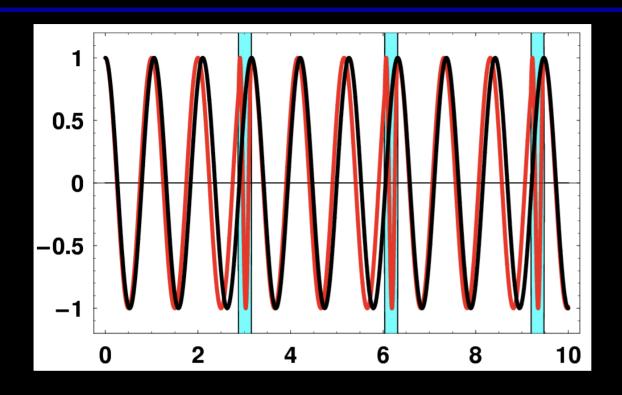
However,





Idea II: Phase Shift Plates







Photon and ALP "remain" in phase!

Can reach higher masses!

Improvement Formula



Before:

$$P_{\gamma \to \text{ALP}} \sim g^2 B^2 \frac{2\sin\left(q\frac{L}{2}\right)}{q}$$

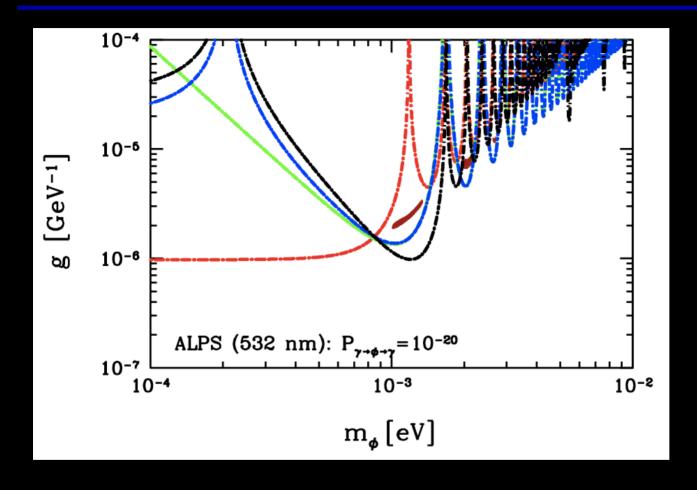
After:

$$P_{\gamma \to \text{ALP}} \sim g^2 B^2 \left[\frac{2 \sin \left(q \frac{L}{2N} \right)}{q} \frac{\sin \left(\frac{N}{2} \left[q \frac{L}{N} + x \right] \right)}{\sin \left(\frac{1}{2} \left[q \frac{L}{N} + x \right] \right)} \right]^2$$
 $\approx g^2 B^2 L^2$

Coherent production over whole length!

Improvement II







Advantages



• Easy to install.

Theorists and "easy" experiments...







Advantages



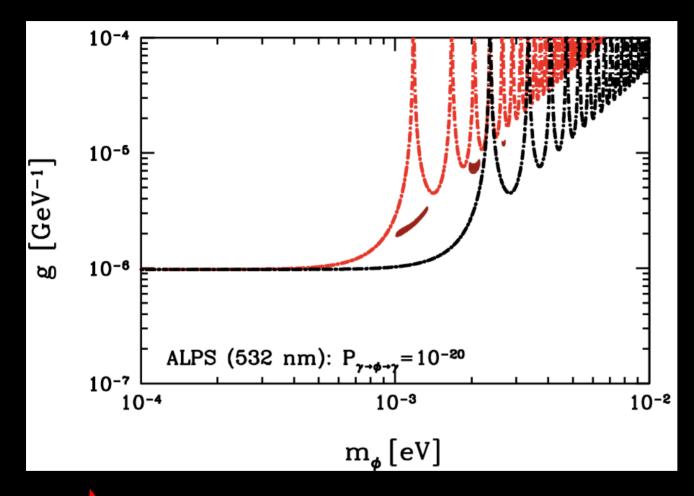
Easy to install.

· Can adapt for different masses.

Allows to scan a mass range!

Improvement II:... scanning







Advantages



· Easy to install.

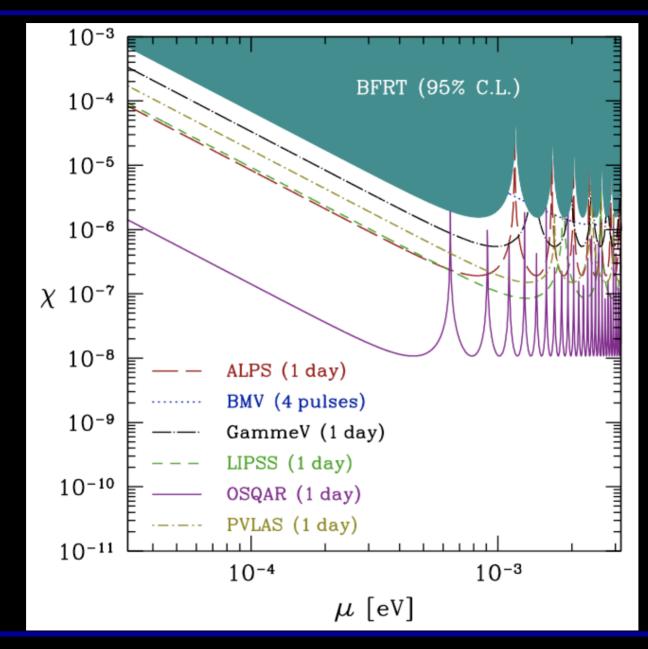
· Can adapt for different masses.

Allows to scan a mass range!

· Works also for Paraphotons!

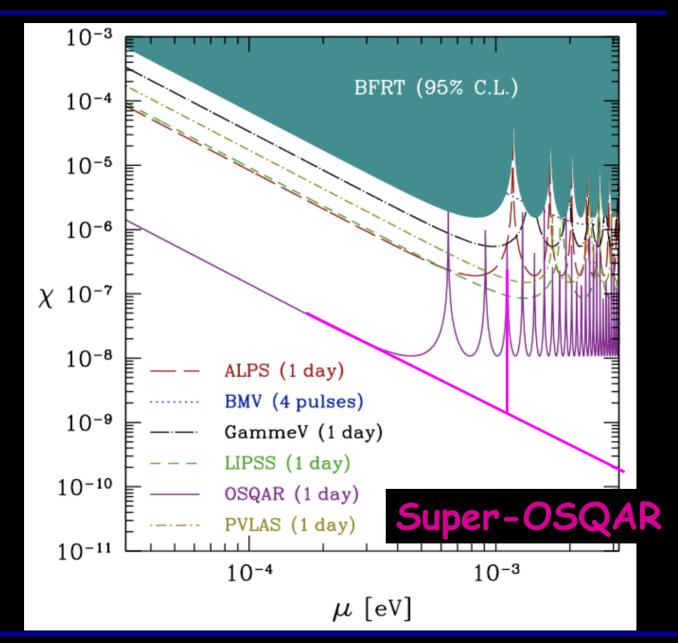
Paraphotons





Paraphotons





Conclusions

Conclusions



- Phase shift plates can improve sensitivity of LSW experiments to higher masses
- Allows to scan mass range
- Works for ALPs and Paraphotons

